Einstein gravity from rational curves in twistor space

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ESI, Vienna, 10/9/2012

Work with Tim Adamo, Freddy Cachazo & David Skinner. Adamo & M arxiv:1203.1026& arxiv:1207.3602, Cachazo & Skinner arxiv:1207.0741 Cachazo, M., Skinner arxiv:1207.4712.

[Cf. also work by Andrew Hodges 1204.1930, Bo Feng & Song He 1207.3220 and 1207.4064, & Mat Bullimore 1207.3940.] Twistors strings \rightsquigarrow remarkable progress for Yang-Mills amplitudes.

What about gravity?

- Twistor-strings \supset Conformal gravity [Berkovits, Witten 2004].
- Twistor-action for conformal gravity [M. 2005].
- But Einstein gravity \subset conformal gravity. \rightarrow two strategies:
 - 1 Try to compute Einstein gravity answer from Berkovits-Witten string (N = 4 SUSY) [Adamo M. 2012].
 - Q Guess full N = 8 formula, at least on momentum space generalizing Hodges' new MHV formula [Cachazo Skinner 2012, + CSM].

In this talk I review and compare the two approaches and further developments.

Definition

For field theory with field g and action S[g] the S-matrix is a sequence of multilinear functionals $C(g_1, \ldots, g_n)$ of n solutions g_i to linearized fields equs. (gives amplitudes for Scattering).

• Construct tree-level S-matrix as action S[g] of solution g to field equs from asymptotic data $g|_{\mathscr{I}} = \sum_{i=1}^{n} \epsilon_i g_i|_{\mathscr{I}}$:

 $\mathcal{C}(g_1,\ldots,g_n) = \text{coeff. of } \prod_i \epsilon_i \text{ in } S_{CG}[g],$

- On 4-dim space-time manifold *M* with metric *g*:
 - Conformal gravity action $S_{CG}[g] := \frac{1}{\kappa^2} \int_M Weyl^2$.
 - Einstein gravity action $S_{EG}[g] := \frac{1}{\kappa^2} \int_M (\text{Scal} + \Lambda) d \text{ vol.}$
- Notation ${\mathcal C}$ as above and Einstein (tree) S-matrix

 $\mathcal{M}(1,\ldots,n) = \text{ Coeff of } \prod_i \epsilon_i \text{ in } S_{EG}[g].$

• Often insert Fourier modes $g_j \sim e^{iP_j \cdot x}$, momentum P_j , $P_j^2 = 0$, so C, \mathcal{M} become functions of P_i .

From conformal gravity to Einstein gravity

Einstein field eqs $R_{ab} = \Lambda g_{ab}$ imply conformal gravity ones $B_{ab} := \Box R_{ab} + \ldots = 0$ so Einstein \subset Conformal gravity.

Proposition (modified Maldacena after Anderson)

The conformal-gravity tree-level S-matrix evaluated on Einstein gravity wave functions with $\Lambda > 0$ gives $\Lambda \times$ Einstein S-matrix.

Proof: (Idea) If *g* is Einstein $R_{ab} = \Lambda g_{ab}$ developed from Einstein data, then:

$$S_{CG}[g] = \int \text{Weyl}^2 = \text{Euler class} + \int \Lambda^2 d\text{vol}$$

whereas

$$S_{EG} = \int \Lambda dvol$$

So also perturbatively $S_{CG} = \Lambda S_{EG} \Rightarrow C = \Lambda M$. Care is needed to get boundary terms right. \Box

N = 4 Super Twistor space and Minkowski space

Spacetime $\mathbb{M} = (\mathbb{R}^{4|8}, g)$ coords $(x^{AA'}, \theta^{aA}) A=0,1,A'=0',1',a=1,..,4$. Twistor space is $\mathbb{PT} = \mathbb{CP}^{3|4}$, homogeneous coords:

$$Z = Z_I = (\lambda_A, \mu^{A'}, \chi^a) \in \mathbb{T} := \mathbb{C}^2 \times \mathbb{C}^2 \times \mathbb{C}^{0|4}, \qquad Z \sim \zeta Z, \zeta \in \mathbb{C}^*$$

 $\mathbb{T} =$ fund. repn of superconformal group SU(2, 2|4).

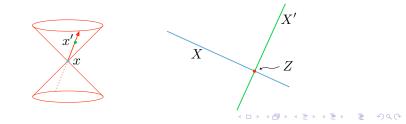
A point $x \in \mathbb{M} \leftrightarrow a$ line $X = \mathbb{CP}^1 \subset \mathbb{PT}$ via incidence relations

$$\mu^{\mathcal{A}'} = -i x^{\mathcal{A}\mathcal{A}'} \lambda_{\mathcal{A}}, \qquad \chi^{\mathcal{A}} = \theta^{\mathcal{A}\mathcal{A}} \lambda_{\mathcal{A}}.$$

Two points x, x' are null separated iff X and X' intersect.

Space-time

Twistor Space



Twistor-strings for conformal gravity Berkovits-Witten 2004

Fields: $Z = Z(\sigma, \bar{\sigma}), Y = Y_{\sigma}(\sigma, \bar{\sigma}) d\sigma, \sigma$ coord on worldsheet Σ : $Z : \Sigma \to \mathbb{PT}, \quad Y \in \Omega^{1,0}(\Sigma) \otimes T^* \mathbb{PT},$

action

$$S[Z, Y, a] = \int_{\Sigma} Y_I \bar{\partial} Z^I, \qquad \bar{\partial} Z = \frac{\partial Z}{\partial \bar{\sigma}} d\bar{\sigma}$$

and field equations $\bar{\partial} Z = 0 = \bar{\partial} Y.$

Data: linear conformal gravity = self-dual \oplus anti-self-dual $F := (f, g) \in H^1(\mathbb{PT}', T \oplus T^*\mathbb{PT}),$

which perturbs action by Vertex operators

$$V_{\mathcal{F}} := V_f + V_g := \int_{\Sigma} f(Z)^I Y_I + g(Z)_I \mathrm{d} Z^I$$

 $V_f \leftrightarrow \text{deformations of } \mathbb{C}\text{-structure on } \mathbb{PT}' \rightsquigarrow \overline{\partial} Z = f(Z),$ V_g gives 'B-field' (noncommutative str?).

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Berkovits-Witten conjecture for conformal gravity

Path integral reduces to integral over space $\mathscr{M}_{d,n}^g$ of holomorphic maps $Z: \Sigma \to \mathbb{PT}$:

$$\mathcal{C}(1,\ldots,n) = \sum_{g,d=0}^{\infty} \int_{\mathcal{M}_{d,n}^{g}} \mathrm{d}\mu_{d} \left\langle V_{F_{1}}(Z(\sigma_{1})) \ldots V_{F_{n}}(Z(\sigma_{n})) \right\rangle_{d} ,$$

g = genus of Σ , d = degree of map, n = # marked points.

- For tree-amplitudes, take g = 0, so $\Sigma \cong \mathbb{CP}^1$.
- Coordinatize Σ with homogeneous coords $\sigma = (\sigma_0, \sigma_1)$.
- So $\mathcal{M}_{d,n} = \text{maps } Z : \mathbb{CP}^1 \to \mathbb{PT}$, degree-d (weight d in σ)

$$Z(\sigma) = \sum_{r=0}^{d} U_r \sigma_0^r \sigma_1^{d-r}, \qquad \mathrm{d}\mu_d = \frac{\prod_r \mathrm{d}^{4|4} U_r}{\mathrm{Vol} \ GL(2)}.$$

Correlator computed from green's function

$$\langle Y_{I}(\sigma) Z_{j}^{J}(\sigma') \rangle_{d} = \frac{\delta_{I}^{J}}{(\sigma \sigma')} \frac{(\xi \sigma')^{d+1}}{(\xi \sigma)^{d+1}}$$

where $(\sigma \sigma') = \sigma_0 \sigma'_1 - \sigma_1 \sigma'_0$ and $\xi \in \Sigma$ is gauge choice.

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Reduction from conformal to Einstein gravity

Linear Einstein SD \oplus ASD fields are given by by

$$H := (h, \tilde{h}) \in H^1(\mathbb{PT}', \mathcal{O}(2) \oplus \mathcal{O}(-2))$$

• Introduce 'infinity twistors' $I_{\alpha\beta}$, $I^{\alpha\beta}$ (fermionic part = 0)

$$I_{lphaeta} = egin{pmatrix} arepsilon^{\mathcal{A}\mathcal{B}} & \mathbf{0} \\ \mathbf{0} & \Lambdaarepsilon_{\mathcal{A}'\mathcal{B}'} \end{pmatrix} \,, \qquad I^{lphaeta} = egin{pmatrix} \Lambdaarepsilon_{\mathcal{A}\mathcal{B}} & \mathbf{0} \\ \mathbf{0} & arepsilon^{\mathcal{A}'\mathcal{B}'} \end{pmatrix} \,.$$

I has rank two when $\Lambda = 0$ ($\Lambda = cosmological const.$) &

$$I^{lphaeta} = rac{1}{2} arepsilon^{lphaeta\gamma\delta} I_{\gamma\delta}\,, \quad I^{lphaeta} I_{eta\gamma} = \Lambda \delta^lpha_\gamma\,,$$

• Gives Poisson structure $\{,\}$ and contact structure τ ,

$$\{h_1, h_2\} := I^{IJ} \partial_I h_1 \partial_J h_2, \qquad \tau = I_{IJ} Z^I \mathrm{d} Z^J,$$

Einstein \subset conformal gravity: $(f^I, g_J) = (I^{IK} \partial_K h, Z^K I_{KJ} \tilde{h})$ so Einstein vertex operators:

$$V_h := \int_{\Sigma} Y \cdot \partial h := \int_{\Sigma} Y_I I^{IJ} \partial_J h, \quad V_{\tilde{h}} = \int_{\Sigma} \tilde{h} \wedge \tau.$$

Einstein amplitudes from twistor-strings

If Berkovits-Witten twistor-string correctly gives conformal gravity amplitudes, then $C = \Lambda M$ gives for Einstein, *k* SD fields:

$$\Lambda \mathcal{M}_{n}^{k} = \int_{\mathcal{M}_{d,n}} d\mu_{d} \left\langle V_{\tilde{h}_{1}} \cdots V_{\tilde{h}_{k}} V_{h_{k+1}} \cdots V_{h_{n}} \right\rangle_{d}$$

$$= \int_{\mathcal{M}_{d,n}} d\mu_{d} \left\langle \tilde{h}_{1}\tau_{1} \cdots \tilde{h}_{k}\tau_{k} Y_{k+1} \cdot \partial h_{k+1} \cdots Y_{n} \cdot \partial h_{n} \right\rangle_{d}$$

- Supersymmetry requires degree of map d = k 1.
- RHS is polynomial degree n in Λ.
- RHS = 0 when $\Lambda = 0$.
- $O(\Lambda)$ part gives Einstein at $\Lambda = 0$.

To check, use momentum eigenstates, momenta $P_j = \lambda_j(ilde{\lambda}_j,\eta_j)$

$$\tilde{h}_{j} = \int_{\mathbb{C}} s_{j} \, \mathrm{d}s_{j} \, \bar{\delta}^{2}(s_{j}\lambda - \lambda_{j}) \, \mathrm{e}^{is_{j}[[\mu \, \tilde{\lambda}_{j}]]} \,, \quad h_{j} = \int_{\mathbb{C}} \frac{\mathrm{d}s_{j}}{s_{j}^{3}} \, \bar{\delta}^{2}(s_{j} - \lambda - \lambda_{j}) \, \mathrm{e}^{is_{j}[[\mu \, \tilde{\lambda}_{j}]]}$$

compute correlator (), integrate & compare to known results.

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compute correlator $\langle \rangle$, integrate & compare to known results.

Lemma

Each $\langle Y_i \cdot \partial h_i \tau_j \rangle$ contraction leads to a power of Λ in the answer. **Proof:** Each such contraction leads to $I^{\alpha\beta}I_{\beta\gamma} = \Lambda \delta^{\alpha}_{\gamma}$. \Box

Degree = 0: maps are just points \rightsquigarrow easy integration over \mathbb{PT} , gives $\Lambda \times$ standard k = 1 3 point amplitude $+\Lambda^2 \times$ new term.

Degree = 1: Maps are now lines \leftrightarrow points in $\mathbb{M}^{4|8}$

 $\mu^{A'} = -i x^{AA'} \lambda_A, \qquad \chi^a = \theta^{aA} \lambda_A, \qquad \lambda_A = \sigma_A.$

fixes 'vol GL(2)', $\mathcal{M}_{1,n} = \mathbb{M}^{4|8} \times (\mathbb{CP}_1)^n$ and $d\mu_1 = d^4x$. Three point k = 2:

- requires one *Y*-contraction $\langle \tilde{h}_1 \tau_1 \tilde{h}_2 \tau_2 Y_3 \cdot \partial h_3 \rangle$.
- Correct answer $O(\Lambda)$ and new term $O(\Lambda^2)$ are obtained

$$\mathcal{C}(1,2,3) = \Lambda \mathcal{M}(1,2,3) = \Lambda \frac{\langle 1 2 \rangle^2}{\langle 1 3 \rangle^2 \langle 2 3 \rangle^2} (1 + \Lambda \Box_{\rho}) \delta^{4|8} \left(\sum_i P_i \right)$$

where $\langle 1 2 \rangle = \lambda_{1A} \lambda_2^A$ also set $[1 2] = \tilde{\lambda}_{1A} \tilde{\lambda}_2^{A'} \oplus \dots \oplus \tilde{\lambda}_{2} \oplus \dots \oplus \tilde{\lambda}_{2}$

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Does $\Lambda = 0$ case give zero? Case: d = 1, k = 2.

$$\int_{\mathscr{M}_{1,n}} d\mu_1 \left\langle \tilde{h}_1 \tau_1 \tilde{h}_2 \tau_2 Y_3 \cdot \partial h_3 \cdots Y_n \cdot \partial h_n \right\rangle_1 = \Lambda \, \mathcal{M}_n^0 = 0$$

• $\langle Y \cdot \partial h_i h_j \rangle = \{h_i, h_j\} = [ij] h_i h_j \text{ at } \Lambda = 0 \text{ (similarly for } \tilde{h}_j) \text{ so each } Y_i \text{ contraction with } h_j, \tilde{h}_j \rightsquigarrow \text{ the factor}$

$$\tilde{\phi}_{j}^{i} := \frac{[ij]\langle \xi j \rangle^{2}}{\langle ij \rangle \langle \xi i \rangle^{2}}, \qquad i \neq j$$

• $\langle Y_i \cdot \partial h_i \tau_j \rangle = O(\Lambda) = 0$, so sum of Y_i contractions \rightsquigarrow factor $-\tilde{\phi}_i^i := \sum_{i \neq i} \frac{[ij] \langle \xi j \rangle^2}{\langle ij \rangle \langle \xi i \rangle^2}$, defines $\tilde{\phi}_j^i$ for i = j

- ξ -independent by momentum conservation $\sum_i \lambda_i \tilde{\lambda}_i = 0$.
- This gives $\prod_{i=3}^{n} (-\tilde{\phi}_{i}^{i})$, but this is generically non-zero!

Resolution: only allow Feynman diagrams for the correlator that are connected trees.

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Proposition

If only connected trees are allowed for the contractions, the conformal gravity amplitude vanishes at $\Lambda = 0$.

Matrix tree theorem: \sum Feynman tree graphs = det $\mathcal{L}_{n-1 n}$ where:

- draw master graph G of all possible (Y_i h_j), (Y_i h̃_j) contractions orienting line from *i* to *j*.
 - 1 n-2 white vertices for $Y_i \cdot \partial h_i$,
 - 2 black vertices for $\tilde{h}_i \tau_i$.
 - **3** white vertices have n 1 outgoing edges, n 3 incoming.
 - 4 black vertices n 2 incoming.
- The weighted Laplacian matrix \mathcal{L} of G has (i, j)-entries

$$\mathcal{L} = \begin{cases} \tilde{\phi}_j^i & i \neq j, n-1, n \\ 0 & i = n-1, n \\ \tilde{\phi}_j^i := -\sum_{j \neq i} \tilde{\phi}_j^i & i = j \end{cases}$$

• Let $\mathcal{L}_{n-1\,n} := \{\mathcal{L} - \text{ rows \& columns } n-1 \text{ and } n\}$

The Hodges matrix arxiv:1204.1930

Hodges obtained a remarkable new formula for the *n*-particley gravitational MHV amplitude as the determinant of an $n - 3 \times n - 3$ minor of an $n \times n$ matrix $\tilde{\Phi}$:

Definition

The Hodges matrix $\tilde{\Phi} := D^{-1} \tilde{\phi} D$ where $D = \text{diag}\{\langle \xi i \rangle^2\}$

Φ
 has co-rank three

$$\sum_{j} ilde{\Phi}^{i}_{j} \lambda_{j \mathcal{A}} \lambda_{j \mathcal{B}} = \mathbf{0} \, .$$

following from momentum conservation.

- $\mathcal{L}_{n-1 n} = \tilde{\phi}_{n-1 n}$ (with n-1 and *n*th row & column removed) as they only differ in the n-1 and *n*th row and column.
- So det $\mathcal{L}_{n-1 n} = \det \tilde{\phi}_{n-1 n} = 0$ as $\tilde{\Phi}$ and $\tilde{\phi}$ have 3-d kernel.

 $\mathcal{C}=0$ at $\Lambda=0$ now follows. \Box

MHV at order Λ , the Hodges formula

Expect gravitational MHV amplitude to be given by

$$\mathcal{M}_{n}^{0} = \frac{1}{\Lambda} \int_{\mathcal{M}_{1,n}} d\mu_{1} \left\langle \tilde{h}_{1}\tau_{1}\tilde{h}_{2}\tau_{2}Y_{3}\cdot\partial h_{3}\cdots Y_{n}\cdot\partial h_{n} \right\rangle_{1} \bigg|_{\Lambda=0}$$

- At $O(\Lambda)$ one Y_i must contract with τ_1 or τ_2 .
- The other n 3 contractions must connect remaining white vertices with one outgoing edge connecting to i, n – 1 or n.
- Matrix tree theorem gives sum of contributions as factor

$$\det \mathcal{L}_{i\,n-1\,n} = \det \tilde{\Phi}_{i\,n-1\,n}$$

multiplied by 3pt amplitude for $\mathcal{M}(i, n-1, n)$.

• This is a version of Hodges' MHV formula.

$$\mathcal{M}(1,\ldots,n) = \mathcal{M}(i,n-1,n) \det \tilde{\Phi}_{i\,n-1\,n}$$

Note: $\tilde{\Phi}$ and Hodges formula have straightforward permutation symmetry and polynomial complexity.

Degree d = k - 1

$$\mathcal{M}_{n}^{k} = \frac{1}{\Lambda} \int_{\mathcal{M}_{k-1,n}} d\mu_{k-1} \left\langle \tilde{h}_{1}\tau_{1}\cdots\tilde{h}_{k}\tau_{k}Y_{k+1}\cdot\partial h_{k+1}\cdots Y_{n}\cdot\partial h_{n} \right\rangle_{k-1} \bigg|_{\Lambda=0}$$

Y_i contractions now give generalized Hodges matrices

$$\tilde{\phi}_{j}^{j} = \begin{cases} \frac{[ij]}{(ij)} \frac{(\xi j)^{k}}{(\xi i)^{k}} & i \neq j \\ -\sum_{l \neq i} \tilde{\phi}_{l}^{j} & i = j \end{cases} \quad (ij) = \sigma_{i\underline{A}}\sigma_{j}^{\underline{A}}$$

• This has co-rank *k* + 1 because relations

$$\sum_{j} \tilde{\phi}_{j}^{j} \sigma_{j A_{1}} \dots \sigma_{j A_{k}} \frac{(\xi \, i)^{k}}{(\xi \, j)^{k}} = \mathbf{0}$$

follow from $\sum_{j} \tilde{\lambda}_{j} \sigma_{jA_{1}} \dots \sigma_{jA_{k-1}} = 0.$

- Matrix-tree thm gives sum of tree contractions as determinant (n - k minor of φ̃) = 0 as co-rank = k + 1.
- At $O(\Lambda)$ with one $\langle Y_i \tau_j \rangle$ contraction, Matrix-tree theorem yields answer as determinant of n k 1 minor of $\tilde{\phi}$.

The Cachazo-Skinner formula for $\mathbb{PT}=\mathbb{CP}^{3|8}$

Theorem (Cachazo, M, Skinner)

The tree-level S-matrix for N = 8 supergravity is given by

$$\mathcal{M}_{n}^{k}(1,\ldots,n) = \int_{\mathcal{M}_{k-1,n}} d\mu_{k-1} \det'(\tilde{\Phi}^{k}) \det'(\Phi^{k}) \prod_{i=1}^{n} \mathrm{D}\sigma_{i} h_{i}(Z(\sigma_{i})),$$

where $\tilde{\Phi}$ is conjugate to $\tilde{\phi}$ as above and

$$\Phi_{i}^{j} = \frac{\langle \lambda(\sigma_{i}) \lambda(\sigma_{j}) \rangle}{(ij)} \qquad i \neq j$$

etc., has rank k - 1 and det' is det of a minor of maximal rank divided by Vandermonde factors in (*i j*).

Proof: Use recursion: shift momenta with complex parameter and show that residues at poles give factorized amplitudes. \Box

Gives full nonlinear (but perturbative) structure of Einstein equations!

Conclusions:

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For Berkovits-Witten twistor-string have good evidence:

- confirmed that twistor-string gives zero at $\Lambda=0$ as required by Maldacena argument.
- obtained Hodges formula for k = 2 (MHV).
- obtained part of $\tilde{\Phi}^k$ in Cachazo-Skinner formula.

But more work required for full understanding.

- N = 8 Cachazo-Skinner formula \rightsquigarrow many new avenues:
 - Is there an *N* = 8 SUGRA twistor-string as well?
 - What is geometric interpretation?
 - Quantization?

Happy Birthday MikE!