Notions of curvature for non-smooth spacetimes

Roland Steinbauer Faculty of Mathematics



Geometry, Analysis, and Physics in Lorentzian Signature BIRS-IMAG Workshop Granda, May 2025

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excellent = austria

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Non-smooth curvature

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Intro & motivation

2 Curvature for rough metrics

- Linear distributional curvature
- Nonlinear distributional curvature

3 Applications

- Impulsive gravitational waves
- Interlude: Causality
- Singularity theorems
- Aside: Synthetic curvature bounds & singularity thms.

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Curvature for non-smooth spacetimes

 Curvature is the essential quantity in GR Einstein equations relate matter/energy to curvature of spacetime

$$\operatorname{Ric} - \frac{1}{2}\operatorname{R} g + \Lambda g = 8 \,\pi \,\mathrm{G}$$

 non-smooth means spacetime metric below g ∈ C² or no spacetime at all (Lorentzian length/metric spaces, causal sets)

Why is this interesting?

- physically relevant models (matched spacetimes, impulsive wave, etc.)
- PDE point-of-view
- *singularities* vs *curvature blow-up CCH* of Penrose
- approaches to Quantum Gravity (no metric, e.g. causal sets)

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Basic geometric properties change if regularity drops

Example 1: Walking on a sphere vs. walking on a cube



It is always shorter to deviate to the right face than to go along the edges.

Example 2: Squeezing the sphere

Convexity fails for metrics of Hölder regularity $g \in C^{1,\alpha}$ $(\alpha < 1)$.



Equator still geodesic but shorter to deviate into hemispheres. [Hartman-Wintner 52]

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Lorentzian causality theory changes if regularity drops

Example 3: Lightcones bubble up

[Chrusciel-Grant 12]



 $g \in C^{0,\alpha}$ ($\alpha < 1$) Non-uniqueness of null geodesics \sim null cone has full measure.

Example 4. The future is not open

[Grant-Kunzinger-Sämann-S 20]



 $g \in C^{0,\alpha} \ (\alpha < 1)$ The blue curve is timelike but reaches $\partial I^+(p)$

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Basic distributional geometry

• recall, distributions on manifolds & distributional tensor fields

 $\mathcal{D}'(M) := \left(\Omega_c^n(M)\right)'$ $\mathcal{D}'_s^r(M) := \left(\mathcal{T}_r^s \otimes_{C^\infty(M)} \Omega_c^n(M)\right)'$ $\cong \mathcal{D}'(M) \otimes_{C^\infty(M)} \mathcal{T}_s^r(M) \cong L_{C^\infty}\left(\mathfrak{X}^*(M)^r, \mathfrak{X}(M)^s; \mathcal{D}'(M)\right)$

distributional metrics

[Marsden 68, Parker 79]

$$g \in \mathcal{D}'_{2}^{0}(M) \cong L_{C^{\infty}}(\mathfrak{X}(M), \mathfrak{X}(M); \mathcal{D}'(M))$$

symmetric and nondeg. $g(X,Y) = 0 \ \forall Y \Rightarrow X = 0 \ (X,Y \in \mathfrak{X}(M))$

- -~g gives no musical isomorphism ${\mathcal D'}_0^1
 i X \mapsto X^{lat} := g(X,.) \in {\mathcal D'}_1^0$
- index, geodesics, etc. of a distributional metric?
- only way to define, inverse, curvature, etc. is via smoothing

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Distributional connections

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- only way to define curvature, etc. is via smoothing
- [LeFloch-Mardare 07]

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- + extend to entire **smooth** tensor algebra
- $+\,$ every \mathcal{D}' -metric has a 'Levi Civita connection'
- used from now on

+ becomes workable if $abla: \mathfrak{X}(M) imes \mathfrak{X}(M) o (L^2)_0^1(M)$

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Curvature from a distributional connection?

- \mathcal{C}^{∞} : Riem : $\mathfrak{X}(M)^3 \to \mathfrak{X}(M)$, $R_{XY}Z := \nabla_{[X,Y]}Z [\nabla_X, \nabla_Y]Z$ (*)
- \mathcal{D}' -connection: ∇ : $\mathfrak{X}(M) \times \mathfrak{X}(M) \to \mathcal{D}'_0^1(M)$

problem: $\nabla_Y Z \in {\mathcal D'}_0^1 \rightsquigarrow \nabla_X \nabla_Y Z$ not defined

• workaround: look for special distributional connections with

 $\nabla : \mathfrak{X}(M) \times \mathfrak{X}(M) \to \mathcal{A}(M) \subseteq \mathcal{D}'_0^1(M)$

such that ∇ can be extended to $\mathcal{A}(M)$ in second slot

$$\nabla: \mathfrak{X}(M) \times \mathcal{A}(M) \to \mathcal{D}'_{0}^{1}(M) \quad (X \in \mathfrak{X}, Y \in \mathcal{A}, \theta \in \Omega^{1})$$
$$\underbrace{\nabla_{X} Y}_{\in \mathcal{D}'_{0}^{1}} := X(\underbrace{Y(\theta)}_{\in \mathcal{D}'}) - \underbrace{\nabla_{X} \theta}_{\in \mathcal{A}}(\underbrace{Y}_{\in \mathcal{A}}) \in \mathcal{D}'(M)$$

• obvious choice $\mathcal{A} = (L^2_{\text{loc}})^1_0$

For such an L^2 -connection, the curvature tensor is defined via (*).

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- want Levi-Civita connection to be $L^2:$ need $\left| \,g \text{ in } H^1_{\text{\tiny loc}} \cap L^\infty_{\text{\tiny loc}} \right.$
- $\bullet\,$ note: $H^1_{\rm loc}\cap L^\infty_{\rm loc}$ is an algebra
- $\bullet\,$ notion of nondegeneracy: $|\det g|\geq C>0$ on compact sets

Definition (Geroch-Traschen 87, LeFloch-Mardare 07, S-Vickers 09)

A distributional metric g is called *gt-regular* if it is of $H^1_{\text{loc}} \cap L^{\infty}_{\text{loc}}$ -regularity and it is uniformly nondegenerate: $\forall K$ cp. $\exists C > 0$: $|\det g| \ge C$ on K.

Geroch-Traschen class is the maximal "reasonable" distributional setting

- + allows to define curvature $\operatorname{Riem}[g]$, $\operatorname{Ric}[g]$, $\operatorname{R}[g]$ in distributions
- + is stable w.r.t. perturbations

Limitations

- $-\,$ Bianchi identities fail \rightarrow energy conservation ?
- $-\dim\left(\operatorname{supp}\left(\operatorname{Riem}[g]\right)\right) \ge n-1 \rightsquigarrow \mathsf{thin \ shells: \ yes, \ strings: \ no!}$

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g be gt-regular & g_{ε} be a (smooth) approximation. When do we have $\operatorname{Riem}[g_{\varepsilon}] \to \operatorname{Riem}[g], \text{ in } \mathcal{D}_3'^1(M)$?

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$$g_{\varepsilon} \to g$$
 in $H^1_{\mathrm{loc}}, \, g_{\varepsilon}^{-1} \to g^{-1} \, \mathrm{in} \, L^{\infty}_{\mathrm{loc}}$

[LeFloch-Mardare 07]

But for smoothings via convolution $g_n^{-1}
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- Existence of approximation with (*)
- if g continuous
- if g not too far from continuous ('stable'): [S-Vickers 09 $\forall K$ cp. there is A^K continuous, such that $\max_{i,j} \operatorname{essup}_{x \in K} |g_{ij}(x) - A^K_{ij}(x)| \le C < \frac{\mu_K}{2n},$ where $\mu_K := \min \operatorname{essinf}_{x \in K} |\lambda^i(x)|, \lambda_i, \dots, \lambda_n$ eigenvalues of a.

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Nonlinear distributions



Nonlinear distributions

Based on *algebras of generalised functions* [Colombeau 84, 85] • differential algebras containing the vector space $\mathcal{D}'(M)$ • display maximal consistency w.r.t. classical analysis; preserve: • the product of C^{∞} functions Lie derivatives of distributions. • regularization of distributions by nets of \mathcal{C}^{∞} -functions • asymptotic estimates in terms of ε (quotient construction) Construction on manifolds $|\mathcal{G}(M) := \mathcal{E}_M(M) / \mathcal{N}(M)|$ [Kunzinger-S 02] $\mathcal{E}_M(M) := \{ (u_{\varepsilon})_{\varepsilon} \in \mathcal{C}^{\infty} : \forall K \forall P \exists l : \sup |Pu_{\varepsilon}(x)| = O(\varepsilon^{-l}) \}$ $x \in K$ $\mathcal{N}(M) := \{ (u_{\varepsilon})_{\varepsilon} \in \mathcal{E}_M(M) : \forall K \forall m : \sup |u_{\varepsilon}(x)| = O(\varepsilon^m) \}$ $x \in K$

fine sheaf of differential algebras w.r.t. $L_X u := [(L_X u_\varepsilon)_\varepsilon]$

• *tensor fields*: fine sheaf of finitely generated, projective $\mathcal{G}(M)$ -modules

$$\begin{aligned} \mathcal{G}_{s}^{r}(M) &:= \mathcal{E}_{s}^{r}(M) / \mathcal{N}_{s}^{r}(M) \\ &\cong L_{\mathcal{C}^{\infty}(M)} \big(\mathcal{G}_{r}^{s}(M), \mathcal{G}(M) \big) \cong \mathcal{G}(M) \otimes_{\mathcal{C}^{\infty}} \mathcal{T}_{s}^{r}(M) \\ &\cong L_{\mathcal{G}(M)} \big(\mathcal{G}_{r}^{s}(M), \mathcal{G}(M) \big) \end{aligned}$$

• embeddings: injective sheaf morphism

 $\iota: \mathcal{T}^r_s(M) \hookrightarrow \mathcal{D}^{\prime r}_s(M) \hookrightarrow \mathcal{G}^r_s(M)$

basically given by chart-wise, component-wise convolution

- generalised metric: $g = [(g_{\varepsilon})_{\varepsilon}] \in \mathcal{G}_2^0(M)$ symm. & det(g) inv. in \mathcal{G} locally represented by sequence of smooth metrics g_{ε} with $|\det(g_{\varepsilon})| \ge \varepsilon^m$ for some m on any compact set.
- musical isomorphism: $\mathcal{G}_0^1(M) \ni X \mapsto X^{\flat} := g(X, .) \in \mathcal{G}_1^0(M)$

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- generalised metric: $g = [(g_{\varepsilon})_{\varepsilon}] \in \mathcal{G}_2^0(M)$ symm. & det(g) inv. in \mathcal{G} locally represented by sequence of smooth metrics g_{ε} with $|\det(g_{\varepsilon})| \ge \varepsilon^m$ for some m on any compact set.
- musical isomorphism: $\mathcal{G}_0^1(M)
 i X \mapsto X^{lat} := g(X, \, . \,) \in \mathcal{G}_1^0(M)$

• *tensor fields*: fine sheaf of finitely generated, projective $\mathcal{G}(M)$ -modules

$$\mathcal{G}_{s}^{r}(M) := \mathcal{E}_{s}^{r}(M) / \mathcal{N}_{s}^{r}(M)$$

$$\cong L_{\mathcal{C}^{\infty}(M)} (\mathcal{G}_{r}^{s}(M), \mathcal{G}(M)) \cong \mathcal{G}(M) \otimes_{\mathcal{C}^{\infty}} \mathcal{T}_{s}^{r}(M)$$

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• embeddings: injective sheaf morphism

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manifold convolution

$$\mathcal{T}_{\varepsilon}(x) := \mathcal{T} \star_{M} \rho_{\varepsilon}(x) := \sum_{i} \chi_{i}(x) \psi_{i}^{*} \Big(\big(\psi_{i*}(\zeta_{i} \cdot \mathcal{T}) \big) * \rho_{\varepsilon} \Big)(x)$$

 ζ_i cut-off functions, ψ_i charts, χ_i partition of unity, ρ_{ε} mollifier on cp. sets, ε small: only finite sum

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Non-smooth curvature

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• every *G*-metric has a *Levi Civita connection*

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$$R_{XY}Z := \nabla_{[X,Y]}Z - [\nabla_X, \nabla_Y]Z \quad \in \mathcal{G}_1^3(M)$$

compatibility with gt-setting

[S-Vickers 09]



technicalities: g stable, ρ admissible to have $|\det(g_{\varepsilon})| \ge \varepsilon^m$

 $\forall j: \int x^{\alpha} \rho_{\varepsilon}(x) \, dx = 0 \quad \text{for all } 1 \leq |\alpha| \leq j \text{ and } \varepsilon \text{ small, and } \forall \eta > 0: \int |\rho_{\varepsilon}(x)| \, dx \leq 1 + \eta \quad \text{for } \varepsilon \text{ small}$

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$$\begin{array}{cccc} H^1_{\mathrm{loc}} \cap L^{\infty}_{\mathrm{loc}} & \ni g & \xrightarrow{\star \rho_{\varepsilon}} & [g_{\varepsilon}] \in \mathcal{G} \\ & & & & \downarrow \mathcal{G} \\ & & & & \downarrow \mathcal{G} \\ & & & & & \downarrow \mathcal{G} \\ & & & & & & \mathrm{Riem}[g] & \xleftarrow{\mathcal{D}' - \mathrm{lim}} & \mathrm{Riem}[g_{\varepsilon}] \end{array}$$

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Intro & motivation

2 Curvature for rough metrics

- Linear distributional curvature
- Nonlinear distributional curvature

3 Applications

- Impulsive gravitational waves
- Interlude: Causality
- Singularity theorems
- Aside: Synthetic curvature bounds & singularity thms.

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- exact models of short but violent burst of gravitational radiation
- non-expanding with Λ

$$\mathrm{d}s^{2} = \frac{2 \,\mathrm{d}\eta \,\mathrm{d}\bar{\eta} - 2 \,\mathrm{d}\mathcal{U} \,\mathrm{d}\mathcal{V} + 2H(\eta,\bar{\eta}) \,\delta(\mathcal{U}) \,\mathrm{d}\mathcal{U}^{2}}{[1 + \frac{1}{6}\Lambda(\eta\bar{\eta} - \mathcal{U}\mathcal{V})]^{2}}$$

geodesic equations ill-defined as distributions

• consistent solution concept in $\mathcal{G}[I,M]$

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BIRS-IMAG, Granada, May 2025 20 / 28

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$$\begin{split} \ddot{U}_{\varepsilon} &= -\left(e + \frac{1}{2} \, \dot{U}_{\varepsilon}^2 \, \tilde{G}_{\varepsilon} - \dot{U}_{\varepsilon} \left(H \, \delta_{\varepsilon} \, U_{\varepsilon}\right)\right) \, \frac{U_{\varepsilon}}{3/\Lambda - U_{\varepsilon}^2 H \delta_{\varepsilon}} \\ & \ddot{Z}_{p\varepsilon} - \frac{1}{2} H_{,p} \, \delta_{\varepsilon} \, \dot{U}_{\varepsilon}^2 = -\left(e + \frac{1}{2} \, \dot{U}_{\varepsilon}^2 \, \tilde{G}_{\varepsilon} - \dot{U}_{\varepsilon} \left(H \, \delta_{\varepsilon} \, U_{\varepsilon}\right)\right) \, \frac{Z_{p\varepsilon}}{3/\Lambda - U_{\varepsilon}^2 H \delta_{\varepsilon}} \\ & \ddot{V}_{\varepsilon} - \frac{1}{2} \, H \, \delta_{\varepsilon}' \, \dot{U}_{\varepsilon}^2 - \delta^{pq} H_{,p} \, \delta_{\varepsilon} \, \dot{Z}_{q}^{\varepsilon} \, \dot{U}_{\varepsilon} = -\left(e + \frac{1}{2} \, \dot{U}_{\varepsilon}^2 \, \tilde{G}_{\varepsilon} - \dot{U}_{\varepsilon} \left(H \, \delta_{\varepsilon} \, U_{\varepsilon}\right)\right) \, \frac{V_{\varepsilon} + H \, \delta_{\varepsilon} U_{\varepsilon}}{3/\Lambda - U_{\varepsilon}^2 H \delta_{\varepsilon}} \\ & \text{where} \qquad \delta_{\varepsilon} = \delta_{\varepsilon} (U_{\varepsilon}(t)) \,, \quad \delta_{\varepsilon}' = \delta_{\varepsilon}' (U_{\varepsilon}(t)) \,, \quad e = 0, \pm 1 \,, \\ & \tilde{G}_{\varepsilon} = \tilde{G}_{\varepsilon} (U_{\varepsilon}(t), Z_{p\varepsilon}(t)) \,, \quad H = H(Z_{p\varepsilon}(t)) \,, \quad \text{and} \quad H_{,p} = H_{,p}(Z_{q\varepsilon}(t)) \end{split}$$

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• recall: below Lipschitz some aspects auf causality break down

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• recall: convexity fails for $g \in C^{1,lpha}$ (lpha < 1) [Hartman-Wintner 52]

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 ω local timelike one-form

Lemma

[Chrusciel-Grant 12, Graf 20, ...]

For $g\in C(M)$ there are $C^\infty\text{-Lorentzian}$ metrics \hat{g}_ε and \check{g}_ε on M with

(i)
$$\check{g}_{\varepsilon} \prec g \prec \hat{g}_{\varepsilon}$$
.

(ii) $\check{g}_{\varepsilon},\,\hat{g}_{\varepsilon}\to g$, and $(\check{g}_{\varepsilon})^{-1},\,(\hat{g}_{\varepsilon})^{-1}\to g^{-1}$ as good as convolution

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Energy conditions and the singularity theorems

• energy conditions/curvature bounds lie at analytical core

 $\operatorname{Ric}(X,X) \ge 0$ for X timelike/null

•
$$g \in C^{1,1} / C^1 / \text{Lip.} \implies \text{Ric}[g] \in L^{\infty}_{\text{loc}} / {\mathcal{D}'}^{(1)} / {\mathcal{D}'}$$

- focusing of geodesics via distributional energy conditions $\langle \operatorname{Ric}[g](X,X), \varphi \rangle \geq 0$ for all non-neg. test *n*-forms φ
- Ine of arguments:
 - ▶ Raychaudhuri/Ricati arguments for CG-regularising sequence \check{g}_{ε}
 - $\operatorname{Ric}[\check{g}_{\varepsilon}] \to \operatorname{Ric}[g]$ only distributional: too weak for positivity
 - ▶ but: $\left(\operatorname{Ric}[g](X,X)\right) \star \rho_{\varepsilon} \geq 0$ for non-neg. ρ_{ε} ▶ $g \in C^1$: $\left|\left(\operatorname{Ric}[g](X,X)\right) \star \rho_{\varepsilon} - \operatorname{Ric}[\check{g}_{\varepsilon}](X,X)\right| \to 0$ loc. uniformly Friedrichs Lemma

plus convergence of geodesics (ODE-theory)

►
$$g \in \text{Lip:} \left| \left(\text{Ric}[g](X, X) \right) \star \rho_{\varepsilon} - \text{Ric}[\check{g}_{\varepsilon}](X, X) \right| \to 0 \text{ only } L^{p}_{\text{loc}} \left(p < \infty \right)$$

and $\text{Ric}[\check{g}_{\varepsilon}] \geq -|\kappa|$

plus worldvolume estimates: see talk by Melanie Graf

Optimal regularity & singularity theorems

work in progress [Kunzinger-Reintjes-S-Vega]

RT-equations

[Reintjes-Temple 20-24]

 $g \in W^{1,p}_{loc} \text{ and } \operatorname{Riem}[g] \in L^p_{loc} \ (n$ $(in a <math>W^{2,p}$ -compatible atlas)

recall

•
$$W^{1,p} \supseteq W^{1,\infty} = \operatorname{Lip}(p < \infty))$$

- $W^{1,p} \subseteq C^{0,\alpha}$ with $\alpha = 1 n/p$ (Morrey's inequality)
- $W^{2,p}$ -change of diff. structure leaves causality invariant (I^{\pm}, J^{\pm})
- Lemma: $g \in W^{1,p}$, $\operatorname{Riem}[g] \in L^p \Longrightarrow$ no causal bubbles
- Hawking & Penrose theorems for $g \in W^{1,p}$, $\operatorname{Riem}[g] \in L^p$ by using the C^1 -results

spacetime results

Penrose, Gannon-Lee, Hawking-Penrose: C¹
 [Graf 20], [Schinnerl-S 22], [Kunzinger-Ohanyan-Schinnerl-S 23]

- Hawking: Lip.
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- Hawking & Penrose: $g \in W^{1,p}$, $\operatorname{Riem}[g] \in L^p$ [work in progress]

[Minguzzi 19]

causal cone structures

synthetic results

- sectional curvature bounds
- Ricci curvature bounds

[Alexander-Graf-Kunzinger-Sämann 22] [Cavaletti-Mondino 22] [Braun-McCann 24]

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causal cone structures

[Minguzzi 19]

- $\bullet\,$ upper semi-cont. distribution of cones on M
- causal core of singularity theorems may be established

Theorem (Causal Penrose)

[Minguzzi 19]

Let (M, C) be a globally hyperbolic closed cone structure w. a non-compact stable Cauchy hypersurface Σ . Then there are no compact future trapped sets and if Σ is non-empty and compact there is an inextendible future null geodesic entirely in $E^+(S)$.

spacetime results

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Outline

Intro & motivation

2 Curvature for rough metrics

- Linear distributional curvature
- Nonlinear distributional curvature

3 Applications

- Impulsive gravitational waves
- Interlude: Causality
- Singularity theorems
- Aside: Synthetic curvature bounds & singularity thms.

- Synthetic approaches: Lorentzian length spaces
 - \blacktriangleright causal space (X,d,\ll,\leq,τ) with τ intrinsic
 - (timelike) sectional curvature bounds via triangle comparison
 - Ricci bds. via optimal transport (RCD-spaces, Lott-Villani, Sturm)
 - ▶ smooth metric measure spacetimes [McCann 20], [Mondino-Suhr 22]
 - ► [Cavaletti-Mondino 22–24] TCD(K,N) and TMCP(K,N) properties

Theorem(TMCP-Hawking)

[Cavaletti-Mondino 22]

Let X be a timelike non-branching, globally hyperbolic LLS with TMCP. Let V be a Borel achronal future timelike complete subset with mean curvature bded above. Then every future timelike geodesic starting in V has a bounded maximal domain of existence.

- synthetic (NEC) [McCann 23]
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- first-order Sobolev calculus on metric measure spacetimes (maximal weak subslope of time functions akin L-modulus of diff.)
 [Beran-Braun-Calisti-Gigli-McCann-Ohanyan-Rott-Sämann 2

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Roland Steinbauer, University of Vienna

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Theorem (Synthetic Hawking) [Alexander-Graf-Kunzinger-Sämann 22]

Let $Y = (a, b) \times_f X$ be a warped product (X metric length space, $f \in C^{\infty}$, non-const.) with positive timelike sectional curvature. Then $a > -\infty$ or $b < \infty$ and hence Y is past/future timelike geodesically incomplete.

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Non-smooth curvature

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