## The wave equation on singular space-times

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- Introduction: Motivation from General Relativity
- ② Digression: Generalized functions & General Relativity
- The wave equation on singular space-times
  - The setting: Foliations etc.
  - O The result: Local existence and uniqueness
  - The method of proof: Higher order energy estimates
- Outlook
- Some references

# Weakly singular space-times I

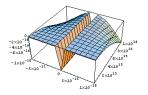
• General Relativity: space-time (M, g)

Einstein equations  $G_{ab}[g] = \kappa T_{ab}$ 

• classically: singularities defined geometrically

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i.e., via incomplete geodesics
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• problem: many reasonable space-times are singular e.g.:



impulsive gravitational waves, shell-crossing singularities, cosmic strings, ...

• often (but not always) the metric is locally bounded but

- below validity of standard differential geometry, i.e.,  $C^{1,1}$
- below largest "reasonable" distributional class, i.e.,  $H^2_{loc} \cap L^{\infty}_{loc}$  [Geroch, Traschen, 87]
- $\rightarrow$  nonlinear distributional geometry (digression)

- physically relevant: not geodesics but tidal forces on extended bodies near singularity
   but unclear how to model!
- idea: study behavior of test fields instead [Clarke, 96] singularities as obstructions to the evolution of test fields rather than as obstructions to extending geodesics
- **so:** study Cauchy problem for the wave equation locally near singularity
- problem:  $\Box_g u = \nabla^a (\nabla_a u) = |\det g|^{-\frac{1}{2}} \partial_a (|\det g|^{-\frac{1}{2}} g^{ab} \partial_b u)$ involves coefficients of low regularity
- results: local unique solvability
  - for shell-crossing singularities [Clarke, 98] (weak solution concept)
  - for cosmic strings [Vickers, Wilson, 00] (in  $\mathcal{G}$ )
  - $\sim$  generalization to "weakly singular" space-times (this talk)

# Distributional differential geometry

[Schwartz, de Rham, Marsden, Parker, ...]

- distributions on manifolds: dual space of compactly supported *n*-forms  $\mathcal{D}'(M) = (\Omega_c^n(M))'$
- distributional sections in vector bundles

$$\mathcal{D}'(M, E) = \left( \Gamma_c(M, E^* \otimes \Omega_c^n(M))' \\ \cong \mathcal{D}'(M) \otimes_{\mathcal{C}^{\infty}(M)} \Gamma(M, E) \\ \cong L_{\mathcal{C}^{\infty}(M)} \left( \Gamma(M, E^*), \mathcal{D}'(M) \right)$$

• extension of classical operations by continuity:

$$L_{\xi}, [,], \wedge, \iota_{\xi}, \{,\} \ldots$$

but with only one  $\mathcal{D}'$ -factor

• distributional pseudo-Riemannian metric

$$g\in {\mathcal D'}_2^0(M)=\{t:\mathfrak{X}(M) imes\mathfrak{X}(M) o {\mathcal D'}(M):t\;\mathcal{C}^\infty(M)-{\sf bilinear}\}$$

symmetric and nondegenerate, i.e.,  $g(\xi,\eta) = 0 \ \forall \eta \in \mathfrak{X}(M) \Rightarrow \xi = 0.$ 

- distributional connection  $D : \mathfrak{X}(M) \times \mathcal{D}'_0^1(M) \to \mathcal{D}'_0^1(M)$  $\mathcal{C}^{\infty}(M)$ -linear in first,  $\mathbb{R}$ -linear in second slot,  $D_{\xi}(f\eta) = fD_{\xi}\eta + \xi(f)\eta$
- problems:
  - can't insert  $\mathcal{D}'\text{-vector}$  fields into  $\mathcal{D}'\text{-metric}$
  - curvature in general undefined:  $R(\xi, \eta) = [D_{\xi}, D_{\eta}] D_{[\xi, \eta]}$  (nonlin.!)
  - notion of **nondegeneracy** is too weak (no point-values  $\rightsquigarrow$  non-local):  $ds^2 = x^2 dx^2$  (classically) degenerate in 0, but nondegenerate in  $\mathcal{D}'$ .
  - index of a distributional metric?
  - geodesics of a distributional metric?

Would need a notion of manifold-valued distributions.

• differential geometry, Einstein equations:

 $g \in \mathcal{D}'(M, T_2^0(M))$  simply not possible!

- positive results in special cases
  - gravitational shock waves [Lichnerowicz, 1950-ies]
  - thin shells of matter, [Israel, 66]
  - ultrarelativistic black holes [Balasin, Nachbagauer, 1990-ies]
  - ...

$$\dim \operatorname{supp}(T_{ab}) = \dim \operatorname{supp}(G_{ab}) = 3$$

- No-go results [Geroch-Traschen, 1987]
  - maximal "reasonable" class is  $g_{ab} \in H^2_{loc} \cap L^\infty_{loc}$ 
    - curvature tensors, in particular  $G_{ab}$  defined in  $\mathcal{D}'$
    - limit consistency, completeness
    - Bianchi identities ( $\rightsquigarrow$  energy conservation) can't be formulated

 $\implies \dim \operatorname{supp}(G_{ab}) \geq 3$ 

• excludes many interesting examples, e.g. cosmic strings

## Nonlinear distributional geometry

• Colombeau tensor fields (special version) usual quotient construction using:

$$(\mathcal{T}_s^r)_M: \ (u_{\varepsilon}) \in (\mathcal{T}_s^r)^{(0,1]}: \ \forall K \subset \subset M \ \forall P \ \exists \ N: \ \sup_{x \in K} \|Pu_{\varepsilon}(x)\|_h = O(\varepsilon^{-N})$$

$$\mathcal{N}_{s}^{r} \quad : \ (u_{\varepsilon}) \in (\mathcal{T}_{s}^{r})^{(0,1]} : \ \forall K \subset \subset M \qquad \forall \ m : \ \sup_{x \in K} \| u_{\varepsilon}(x) \|_{h} = O(\varepsilon^{m})$$

$$\mathcal{G}_{s}^{r}(M) := (\mathcal{T}_{s}^{r})_{M} / \mathcal{N}_{s}^{r} \qquad \mathcal{G}(M) := \mathcal{G}_{0}^{0}(M)$$

- G(\_) is a fine sheaf of differential algebras w.r.t. L<sub>X</sub> u := [(L<sub>X</sub> u<sub>ε</sub>)<sub>ε</sub>]
   G<sup>r</sup><sub>s</sub>(\_) is a fine sheaf of C<sup>∞</sup>(M)-modules and of G(M)-modules; finitely gen., projective
- characterizations:

$$\mathcal{G}_{s}^{r}(M) \cong \mathcal{G}(M) \otimes \mathcal{T}_{s}^{r}(M) \cong L_{\mathcal{C}^{\infty}(M)}\Big(\mathcal{T}_{r}^{s}(M), \mathcal{G}(M)\Big)$$
$$\cong L_{\mathcal{G}(M)}\Big(\mathcal{G}_{r}^{s}(M), \mathcal{G}(M)\Big)$$

• For  $g \in \mathcal{G}_2^0(M)$  TFAE

- (i)  $g: \mathcal{G}_0^1(M) \times \mathcal{G}_0^1(M) \to \mathcal{G}(M)$  symm. & det g invertible in  $\mathcal{G}(M)$
- (ii)  $\forall \tilde{x}: g(\tilde{x}): \widetilde{\mathbb{K}}^n \times \widetilde{\mathbb{K}}^n \to \widetilde{\mathbb{K}}$  symmetric and nondegenerate
- (iii) On every rel. cp. chart there exists a representative that is a classical pseudo-Riemannian metric and det g is invertible in  $\mathcal{G}(M)$
- If there exists a representative with constant index, g is called **generalized pseudo-Riemannian metric**.

results:

• index of is well-defined

(proof by finite-dimensional perturbation theory)

- inverse metric  $g^{-1}$  is a well-defined element of  $\mathcal{G}^2_0(M)$
- $\xi \mapsto g(\xi, .)$  induces a  $\mathcal{G}(M)$ -linear isomorphism:  $\mathcal{G}_0^1(M) \to \mathcal{G}_1^0(M)$  $\rightsquigarrow$  raising/lowering of indices

• Generalized connection:

$$\hat{D}:\mathcal{G}_0^1(M) imes \mathcal{G}_0^1(M) o \mathcal{G}_0^1(M)$$

 $\mathcal{G}(M)$ -linear in first,  $\mathbb{\tilde{R}}$ -linear in second slot,  $\hat{D}_{\xi}(f\eta) = f\hat{D}_{\xi}\eta + \xi(f)\eta$ 

- **Fundamental Lemma:** For every generalized metric there exists a unique metric and torsion-free connection, the *generalized Levi-Civita connection*.
- curvature can be defined

$$\hat{R}_{\xi,\eta}\zeta := \hat{D}_{[\xi,\eta]}\zeta - [\hat{D}_{\xi},\hat{D}_{\eta}]\zeta$$

Einstein equations can be formulated

compatibility with

- classical theory:  $\hat{g} \approx_k g \Rightarrow$  gen. curvature  $\approx_{k-2}$  class. curvature
- Geroch-Traschen approach (recent work)

• **The Result:** Local existence and uniqueness Theorem for the Cauchy problem for the scalar wave equation for a large class of "weakly singular" (essentially locally bounded) space-time metrics in the (special) Colomebau algebra.

### Class of metrics:

precise definition needs some geometric preparation

- $\bullet\,$  formulated entirely in  ${\cal G}$
- many weakly singular metrics can be modelled within this class
  - embedding via convolution with standard mollifier
  - physically motivated modelling
- e.g.: impulsive pp-waves (in Rosen form), expanding spherical impulsive waves, cosmic strings, ...

the metric: g<sub>ab</sub> ∈ G<sub>2</sub><sup>0</sup>(M) a generalized Lorentz metric
 ⇒ ∀V ⊆ M rel. cp. ∃ repr. (g<sup>ε</sup><sub>ab</sub>)<sub>ε</sub> which is a classical L-metric and
 ∀K ⊂⊂ U∃N : inf<sub>K</sub>|| det g<sub>ε</sub>|| ≥ ε<sup>m</sup>

use exclusively such representatives from now on

- the geometry:  $p \in U \subseteq M$  open & rel. cp.  $t: U \to \mathbb{R}$  smooth, t(p) = 0,  $\sigma := dt \neq 0$  on Uassume  $\exists M_0 > 0: g_{\varepsilon}^{-1}(dt, dt) \leq 1/M_0^2$  $\Longrightarrow \Sigma_{\tau} := \{q \in U: t(q) = \tau\}$  "uniformly spacelike" hypersurfaces
- constructions: normal vector  $\xi_{\varepsilon}^{a} := g_{\varepsilon}^{ab} \sigma_{b}$ ; norm  $V_{\varepsilon}^{2} := -g_{\varepsilon}(\xi_{\varepsilon}, \xi_{\varepsilon})$ corresponding normalized versions:  $\hat{\xi}_{\varepsilon}^{a} = \xi_{\varepsilon}^{a}/V_{\varepsilon}$ ,  $\hat{\sigma}_{a}^{\varepsilon} := \sigma_{a}/V_{\varepsilon}$

note: only  $\sigma$  independent of  $\varepsilon$ !

generalized Riemann metric:  $e_{ab}^{\varepsilon} := g_{ab}^{\varepsilon} + 2\hat{\sigma}_{a}^{\varepsilon}\hat{\sigma}_{b}^{\varepsilon}$ 

We define the following class of **weakly singular metrics** (A)  $\forall K \subset U \ \forall k \ \forall \eta_1, \ldots, \eta_k \in \mathfrak{X}(M)$ •  $\sup_{\kappa} ||L\eta_1 \dots L\eta_k g_{\varepsilon}|| = O(\varepsilon^{-k})$ •  $\sup_{\kappa} ||L\eta_1 \dots L\eta_k g_{\varepsilon}^{-1}|| = O(\varepsilon^{-k})$ in particular  $g_{\varepsilon}$ ,  $g_{\varepsilon}^{-1}$  locally uniformly bounded  $\Rightarrow 1/M_0 \leq \sqrt{-g_{\varepsilon}(\xi_{\varepsilon},\xi_{\varepsilon})} = V_{\varepsilon} \leq M_0$ (B)  $\forall K \subset \subset U$ :  $\sup_{K} ||\nabla_{g^{\varepsilon}}\xi_{\varepsilon}|| = O(1)$  $\Rightarrow ||L_{\varepsilon}g_{\varepsilon}||_{e_{\varepsilon}} = O(1), ||\nabla_{g^{\varepsilon}}e_{\varepsilon}|| = O(1) = ||\nabla_{g^{\varepsilon}}e_{\varepsilon}^{-1}||_{e_{\varepsilon}}$ (C) for each  $\varepsilon : \Sigma := \Sigma_0$  past cp., spacelike hypersurface s.t.  $\partial J_{\varepsilon}^+(\Sigma) = \Sigma$ and  $\exists A \neq \emptyset$ , open  $A \subseteq \bigcap_{c} J_{c}^{+}(\Sigma)$  $\Rightarrow$  existence of classical solutions on common domain all norms || || derived from some classical Riemannian background metric

### Formulation of the result

• For given weakly singular g and  $v, w \in \mathcal{G}(\Sigma)$  we consider the i.v.p.

$$\Box_g u = 0, \qquad u|_{\Sigma} = v, \quad L_{\hat{\xi}}|_{\Sigma} u = w \tag{\Box}$$

and ask for solutions in  $\mathcal{G}(U)$ 

- **Theorem:** Let (M, g) be a generalized space-time with a weakly singular metric (i.e., (A)–(C) hold) and let  $v, w \in \mathcal{G}(\Sigma)$ . Then for any  $p \in \Sigma$  there exists an open neighborhood V where  $(\Box)$  has a unique solution in  $\mathcal{G}(V)$ .
  - In coordinates  $(U, \{t, x^1, x^2, x^3\})$   $g_{\varepsilon} = -V_{\varepsilon}^2 dt^2 + h_{ij}^{\varepsilon} (dx^i - N_{\varepsilon}^i dt) \otimes (dx^j - N_{\varepsilon}^j dt)$  $V_{\varepsilon}, \ \partial V_{\varepsilon}, \ h_{ij}^{\varepsilon} = O(1), \qquad \partial^{\alpha} V_{\varepsilon}, \ \partial^{\alpha} h_{ij}^{\varepsilon}, \ \partial^{\alpha} N_{\varepsilon}^i = O(\varepsilon^{|\alpha|})$

$$\Box_{g^{\varepsilon}} u_{\varepsilon} = |\det g_{\varepsilon}|^{-\frac{1}{2}} \partial_{a} (|\det g_{\varepsilon}|^{-\frac{1}{2}} g_{\varepsilon}^{ab} \partial_{b}) u_{\varepsilon} = f_{\varepsilon} \in \mathcal{N}(U)$$
$$u_{\varepsilon} (t = 0, x^{i}) = v_{\varepsilon} (x^{i}), \quad L_{\hat{\xi}_{\varepsilon}} u_{\varepsilon} (t = 0, x^{i}) = w_{\varepsilon} (x^{i}) \qquad (\Box_{\varepsilon})$$

# General strategy of proof

key ingredient: higher order energy estimates

- use of  $\varepsilon$ -dependent Sobolev norms
- use of  $\varepsilon$ -dependent energy-momentum tensors
- switch back to sup-norms

**step 1:** existence of classical solutions  $u_{\varepsilon}$  for fixed  $\varepsilon$  (by (C))

 $\rightsquigarrow$  candidate for  $\mathcal G\text{-solution}$ 

**step 2:** existence of  $\mathcal{G}$ -solutions: moderate data  $\rightsquigarrow$  moderate initial energies  $\rightsquigarrow$  moderate energies for all times  $\rightsquigarrow (u_{\varepsilon})_{\varepsilon} \in \mathcal{E}_{M}(V)$ **step 3:** uniqueness—part 1: independence of the choice of the representatives of v, wnegligible data  $\rightsquigarrow \qquad [\dots ] \qquad \rightsquigarrow (u_{\varepsilon})_{\varepsilon} \in \mathcal{N}(V)$ **step 2:** uniqueness—part 2: independence of the choice of the representative of g

### Sobolev norms and energies I

• use the R-metric  $e_{ab}^{\varepsilon} := g_{ab}^{\varepsilon} + 2\hat{\sigma}_{a}^{\varepsilon}\hat{\sigma}_{b}^{\varepsilon}$  to define "pointwise" norm  $|\nabla_{\varepsilon}^{(j)}u|^2 := ||\nabla_{p_1}^{\varepsilon}\dots\nabla_{p_i}^{\varepsilon}u||_{e_{\varepsilon}}^2$  $= e_{\varepsilon}^{p_1q_1} \dots e_{\varepsilon}^{p_jq_j} (\nabla_{p_1}^{\varepsilon} \dots \nabla_{p_i}^{\varepsilon} u) (\nabla_{q_1}^{\varepsilon} \dots \nabla_{q_i}^{\varepsilon} u)$ • 4D-Sobolev norms on  $\Omega_{\tau}$  w.r.t.  $\nabla_{g^{\varepsilon}}$  resp.  $\partial$  $\nabla \|u\|_{\Omega_{ au,\varepsilon}}^k := \left(\sum_{j=0}^k \int_{\Omega_{ au}} |
abla_{\varepsilon}^{(j)}(u)|^2 \mu^{\varepsilon}
ight)^{rac{1}{2}}$  $\partial \|u\|_{\Omega_{\tau},\varepsilon}^{k} := \Big(\sum_{p_{1},\dots,p_{i}} \int_{\Omega_{\tau}} |\partial_{p_{1}}\dots\partial_{p_{j}}u|^{2}\mu^{\varepsilon}\Big)^{\frac{1}{2}}$ 0<i<  $S_{\tau}$  $\Sigma_{\tau}$  $S_{\Omega,\tau}$ Ω, S Σ  $x^i$ р (Osaka 2008) February 2008 16 / 23

# Sobolev norms and Energies II

• **3D-Sobolev norms on** 
$$S_{\tau}$$
 w.r.t.  $\nabla_{g^{\varepsilon}}$  resp.  $\partial$   
 $\nabla \|u\|_{S_{\tau},\varepsilon}^{k} := \left(\sum_{j=0}^{k} \int_{S_{\tau}} |\nabla_{\varepsilon}^{(j)}(u)|^{2} \mu_{\tau}^{\varepsilon}\right)^{\frac{1}{2}}$   
 $\partial \|u\|_{S_{\tau},\varepsilon}^{k} := \left(\sum_{\substack{p_{1},\dots,p_{j}\\0 \le j \le k}} \int_{S_{\tau}} |\partial_{p_{1}}\dots\partial_{p_{j}}u|^{2} \mu_{\tau}^{\varepsilon}\right)^{\frac{1}{2}}$ 

• Energy momentum tensors (k > 0)

$$\begin{aligned} T_{\varepsilon}^{ab,0}(u) &:= -\frac{1}{2}g_{\varepsilon}^{ab}u^2 \\ T_{\varepsilon}^{ab,k}(u) &:= (g_{\varepsilon}^{ac}g_{\varepsilon}^{bd} - \frac{1}{2}g_{\varepsilon}^{ab}g_{\varepsilon}^{cd})e_{\varepsilon}^{p_1q_1}\dots e_{\varepsilon}^{p_{k-1}q_{k-1}} \\ (\nabla_{\varepsilon}^{\varepsilon}\nabla_{p_1}^{\varepsilon}\dots\nabla_{p_{k-1}}^{\varepsilon}u)(\nabla_{d}^{\varepsilon}\nabla_{q_1}^{\varepsilon}\dots\nabla_{q_{k-1}}^{\varepsilon}u) \end{aligned}$$

• Energies  $(k \ge 0)$  $E_{\tau,\varepsilon}^k(u) := \sum_{j=0}^k \int_{S_{\tau}} T_{\varepsilon}^{ab,j}(u) \widehat{\xi}_a^{\varepsilon} \widehat{\xi}_b^{\varepsilon} \widehat{\mu}_{\tau}^{\varepsilon}$ 

# Lemma 1: Equivalency of norms and energies

(i)  $(\nabla \parallel \parallel_{S_{\tau},\varepsilon} \text{ vs. energies})$ 

There exist constants A, A' such that

$$A'(^{\nabla} \|u\|_{\mathcal{S}_{\tau},\varepsilon}^k)^2 \leq E_{\tau,\varepsilon}^k(u) \leq A(^{\nabla} \|u\|_{\mathcal{S}_{\tau},\varepsilon}^k)^2$$

(ii)  $(\nabla \| \|_{S_{\tau},\varepsilon} \text{ vs. } \partial \| \|_{S_{\tau},\varepsilon})$ 

For each k there exist positive constants  $B_k, B'_k$  such that

$$(\nabla \|u\|_{S_{\tau,\varepsilon}}^{k})^{2} \leq B'_{k} \sum_{j=1}^{k} \frac{1}{\varepsilon^{2(k-j)}} (\partial \|u\|_{S_{\tau,\varepsilon}}^{j})^{2}$$

$$(\partial \|u\|_{S_{\tau,\varepsilon}}^{k})^{2} \leq B_{k} \sum_{j=1}^{k} \frac{1}{\varepsilon^{2(k-j)}} (\nabla \|u\|_{S_{\tau,\varepsilon}}^{j})^{2}$$

(iii) (basic energy estimate)  $\forall \varepsilon, \tau$  the dominant energy condition gives:

$$\mathsf{E}^{k}_{\tau,\varepsilon}(u) \leq \mathsf{E}^{k}_{\tau=0,\varepsilon}(u) + \sum_{j=0}^{k} \int_{\Omega_{\tau}} \left( \xi_{b} \nabla^{\varepsilon}_{\mathsf{a}} T^{\mathsf{ab},j}_{\varepsilon}(u) + T^{\mathsf{ab},j}_{\varepsilon}(u) \nabla^{\varepsilon}_{\mathsf{a}} \xi_{b} \right) \mu_{\varepsilon}.$$

# Proposition 1: Energy estimates

Let  $u_{\varepsilon}$  be a solution of  $(\Box_{\varepsilon})$ . Then  $\forall k \exists C'_k, C''_k, C'''_k$  s.t.  $\forall 0 \leq \tau \leq \gamma$ (i)  $E_{\tau,\varepsilon}^{k}(u_{\varepsilon}) \leq E_{0,\varepsilon}^{k}(u_{\varepsilon}) + C_{k}^{\prime}(\nabla \|f_{\varepsilon}\|_{\Omega_{\tau},\varepsilon}^{k-1})^{2} + C_{k}^{\prime\prime\prime}\int_{\zeta-0}^{\tau} E_{\zeta,\varepsilon}^{k}(u_{\varepsilon})d\zeta$  $+ C_k'' \sum_{i=1}^{k-1} \frac{1}{\varepsilon^{2(1+k-j)}} \int_{\zeta=0}^{\tau} E_{\zeta,\varepsilon}^j(u_{\varepsilon}) d\zeta$ (ii)  $E_{ au,\varepsilon}^k(u_{\varepsilon}) \leq \left(E_{0,\varepsilon}^k(u_{\varepsilon}) + C_k'(\nabla \|f_{\varepsilon}\|_{\Omega_{ au,\varepsilon}}^{k-1})^2\right)$  $+ C_k'' \sum_{i=1}^{k-1} \frac{1}{\varepsilon^{2(1+k-j)}} \int_{\varepsilon}^{t} E_{\zeta,\varepsilon}^j(u_{\varepsilon}) d\zeta \Big) e^{C_k'''\tau}$ 

(iii) If the initial energy  $(E_{0,\,\varepsilon}^k(u_{\varepsilon}))_{\varepsilon}$  is moderate [negligible] then

$$\sup_{0\leq\tau\leq\gamma}(\textit{E}^{\textit{k}}_{\tau,\,\varepsilon}(\textit{u}_{\varepsilon}))_{\varepsilon}$$

is moderate [negligible].

### (i) (Bounds on initial energies from initial data)

Let  $u_{\varepsilon}$  be a solution of  $(\Box_{\varepsilon})$ . If  $(v_{\varepsilon})_{\varepsilon}$ ,  $(w_{\varepsilon})_{\varepsilon}$  are moderate resp. negligible, then the initial energies  $(E_{0,\varepsilon}^{k}(u_{\varepsilon}))_{\varepsilon}$ , for each  $k \ge 0$ , are moderate resp. negligible nets of real numbers. [estimate *t*-derivatives of the data using the wave equation  $\partial_{t}^{2}u_{\varepsilon} = -V_{\varepsilon}^{2}\left(f_{\varepsilon} + \frac{2}{V_{\varepsilon}^{2}}N^{i}\partial_{t}\partial_{i}u_{\varepsilon} - \left(h_{\varepsilon}^{ij} - \frac{1}{V_{\varepsilon}^{2}}N_{\varepsilon}^{i}N_{\varepsilon}^{j}\right)\partial_{i}\partial_{j}u_{\varepsilon} + g_{\varepsilon}^{ab}\Gamma[\mathbf{g}_{\varepsilon}]^{c}{}_{ab}\frac{\partial u_{\varepsilon}}{\partial x^{c}}\right)]$ 

(ii) (Bounds on solutions from bounds on energies) For m > 3/2 an integer, there exists a constant K and number N such that for all  $u \in C^{\infty}(\Omega_{\tau})$  and for all  $\zeta \in [0, \tau]$  we have

$$\sup_{x\in\Omega_{\tau}}|\partial_{x^{i_1}}\cdots\partial_{x^{i_l}}u(x)|\leq K\varepsilon^{-N}\sup_{0\leq\zeta\leq\tau}E_{\zeta,\varepsilon}^{m+l}(u).$$

[uses Sobolev embedding plus Lemma 1]

(i) Let  $\widehat{g}_{\varepsilon}$  be another representative and consider

$$\begin{array}{rcl} \Box_{\hat{g}^{\varepsilon}} \hat{u}_{\varepsilon} &=& f_{\varepsilon} \\ \hat{u}_{\varepsilon}(t=0,x^{i}) &=& v_{\varepsilon}(x^{i}) \\ \partial_{t} \hat{u}_{\varepsilon}(t=0,x^{i}) &=& w_{\varepsilon}(x^{i}) \end{array}$$

(ii) The difference  $\widetilde{u}_arepsilon:=u_arepsilon-\widehat{u}_arepsilon$  solves

$$\Box_{\hat{g}^{\varepsilon}}\widetilde{u}_{\varepsilon}=f_{\varepsilon}-\Box_{\hat{g}^{\varepsilon}}u_{\varepsilon}, \text{ vanishing data}$$

(iii)  $f_{\varepsilon} - \Box_{\hat{g}^{\varepsilon}} u_{\varepsilon} = (f_{\varepsilon} - \Box_{g^{\varepsilon}} u_{\varepsilon}) + (\Box_{g^{\varepsilon}} u_{\varepsilon} - \Box_{\hat{g}^{\varepsilon}} u_{\varepsilon}) = \Box_{g^{\varepsilon}} u_{\varepsilon} - \Box_{\hat{g}^{\varepsilon}} u_{\varepsilon}$ is negligible by well-definedness of  $\Box_{g}$ .

- more general classes of metrics: log-type growth in  $\varepsilon$  replacing O(1)...Hölder-Zygmund classes
- generalization of Hadamard parametrix constructions
- generalized singularity theorems

(extending geodesics in generalized sense)

- go non-linear: Einstein equations
- connections to cosmic censorship hypothesis

# Some references

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