The classical singularity theorems of GR under optimal regularity conditions

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Introduction: The singularity theorems of GR & regularity issues

2 The Hawking-Penrose Theorem: From classical to $C^{1,1}$

3 C^{1,1}-genericity & SEC force causal geodesics to stop maximising

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Singularity Theorems in GR

- singularities occur in exact solutions; high degree of symmetries
- singularities as obstruction to extend causal geodesics [Penrose, 65]

Theorem (Pattern singularity theorem [Senovilla 98])

A spacetime is causal geodesically incomplete if we have (i) an energy/curvature condition, (iii) an initial or boundary (ii) a causality condition, and condition

- (iii) initial condition \rightsquigarrow causal geodesics start focussing
- (i) energy condition \rightsquigarrow focusing goes on \rightsquigarrow focal point
- (ii) causality condition \rightsquigarrow no focal points
- way out: one causal geodesic has to be incomplete

The issue of regularity

Theorem (Pattern singularity theorem [Senovilla 98])

A C²-spacetime ¹ is causal geodesically incomplete if we have
(i) an energy/curvature condition, (iii) an initial or boundary

(ii) a causality condition, and

condition

- C² is too much to ask for: Realistic models (stars, matched spacetimes) involve jumps in matter variables → g ∈ C^{1,1}.
- Theorem allows (i)–(iii) plus completeness for $C^{1,1}$.
- But $C^{1,1}$ -spacetimes are not 'singular' (curvature bd., geodesics ok).
- Below $C^{1,1}$: unbounded curvature, non-unique geodesics: singular.

Hence $C^{1,1}$ is the natural regularity for the singularity theorems.

 $^{1}(M,g)$ with M smooth $g \in C^{2}$

Low $(=C^{1,1})$ regularity: Problems & Solutions

Problems:

- Curvature tensor only $L^\infty \rightsquigarrow$ no Jacobi fields, conjugate/focal points
- No second variation of arclength
- exp_p not a local diffeomorphism.

However:

- exp_p bi-Lipschitz homeomorphism and ∃ convex neighbourhoods, Gauss Lemma holds [Minguzzi 14], [Kunzinger, S, Stojković 14]
- Bulk of causality theory remains valid [Chruściel, Grant 12] [Minguzzi 15] [Kunzinger, S, Stojković, Vickers 14], [Sämann 16]
- The Hawking singularity theorem (big bang) holds in C^{1,1} [Kunzinger, S, Stojković, Vickers 15]
- The Penrose singularity theorem (black hole) holds in C^{1,1} [Kunzinger, S, Vickers 15]

Strategies in low regularity

(1) CG-regularization of the metric adapted to causal structure



Sandwich null cones of $g \in C^0$ between null cones of two approximating families of smooth metrics: $\check{\mathbf{g}}_{\varepsilon} \prec \mathbf{g} \prec \hat{\mathbf{g}}_{\varepsilon}$

[Chruściel, Grant 12]

(2) Use replacement for strong energy condition

Lemma (timelike case)

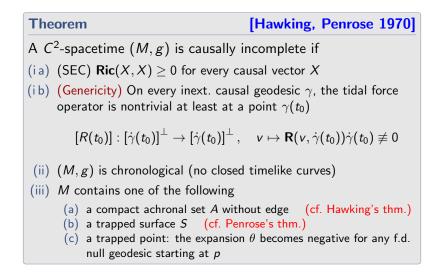
Let (M, \mathbf{g}) be a $\mathcal{C}^{1,1}$ -spacetime satisfying the energy condition Ric $[\mathbf{g}](X, X) \ge 0$ a.e. for all timelike local \mathcal{C}^{∞} -vector fields X. Then for all $K \subset \subset M \quad \forall C > 0 \quad \forall \delta > 0 \quad \forall \kappa < 0 \quad \forall \varepsilon$ small Ric $[\check{\mathbf{g}}_{\varepsilon}](X, X) > -\delta \quad \forall X \in TM|_{\mathcal{K}} : \check{\mathbf{g}}_{\varepsilon}(X, X) \le \kappa, \quad ||X||_{h} \le C.$

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The Hawking-Penrose Theorem



Comments on the classical proof

Proof rests on

The Hawking-Penrose Lemma [Hawking, Penrose 1970] [Graf 2016]

A $C^2C^{1,1}$ -spacetime (M,g) is causally incomplete if

(L1) M is chronological

(L2) Every complete causal geodesic contains a pair of conjugate ptsis not (globally) maximising

(L3) There is a trapped set (S achronal, $E^+(S) := J^+(S) \setminus I^+(S)$ cp.)

Good news: The H-P Lemma continues to hold in $C^{1,1}$ (causality) Left to do: Show that

- appropriate version of the initial conditions \Rightarrow (L3) (causality)
- appropriate version of genericity and SEC \Rightarrow (L2) (analysis, here!)

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The $\mathcal{C}^{1,1}$ -genericity condition

 $\begin{array}{l} \hline \textbf{Definition} \ (\mathcal{C}^{1,1}\text{-genericity condition}) \\ \hline \textbf{Genericity holds along a causal geodesic } \gamma \text{ of a } \mathcal{C}^{1,1}\text{-metric } g \text{ if} \\ \textbf{near some } \gamma(t_0) \text{ there are continuous vector fields } X, V \text{ with} \\ X(\gamma(t)) = \dot{\gamma}(t), V(\gamma(t)) \in \dot{\gamma}(t)^{\perp} \text{ such that} \\ & \langle \textbf{R}(V,X)X,V \rangle > c. \end{array}$

- Equivalent to the usual condition for $g \in C^2$
- Survives approximation process (Friedrichs lemma): If $\gamma_{arepsilon} o \gamma$ in \mathcal{C}^1

$$R[g_{\varepsilon}](t) > \operatorname{diag}(c, -C, \dots, -C) \text{ on } [t_0 - r, t_0 + r]$$
(1)

where $R[g_{\varepsilon}](t) := R[g_{\varepsilon}](.,\dot{\gamma}_{\varepsilon}(t))\dot{\gamma}_{\varepsilon}(t): \dot{\gamma}_{\varepsilon}(t)^{\perp}
ightarrow \dot{\gamma}_{\varepsilon}(t)^{\perp}$

• to be fed into a matrix Riccati comparison argument later on...

Raychaudhuri argument (timelike case)

- γ tl. geodesic in approximating C^{∞} -spacetime, no conjugate pts.
- A (unique) Jacobi tensor with A(-T) = 0 and $A(t_0 = 0) = id$
- $B := A' A^{-1}$, expansion $\theta = tr(B)$ satisfies **Raychaudhuri eq.**:

 $\dot{\theta} = -\operatorname{Ric}(\dot{\gamma}, \dot{\gamma}) - \operatorname{tr}(\sigma^2) - (1/d) \theta^2$

- 'old' (direct) argument: SEC $\Rightarrow \dot{\theta} \leq \delta \frac{1}{d}\dot{\theta}^2$; i.c. $\Rightarrow \theta(0) < b < 0$ \Rightarrow upper bd. on first conj. pt in terms of *b* (scalar Riccati comp.)
- 'reverse' Raychaudhuri: no conj. pts. \Rightarrow $|\theta|$ small initially

Boxing lemma

For T > 0 there is $\delta(T) > 0$ such that: If γ is has no conjugate points on [-T, T] then $\sup_{t \to 0} |\theta(t)| < \frac{4d}{2t}$

$$t \in \left[-\frac{T}{2}, \frac{T}{2}\right] \quad |t| \in \left[-\frac{T}{2}, \frac{T}{2}\right]$$

provided that $\operatorname{Ric}(\dot{\gamma},\dot{\gamma}) \geq -\delta$ on [-T, T].

Matrix Riccati comparison argument

- $B := A' A^{-1}$ satisfies a matrix Riccati eq.: $\dot{B} + B^2 + R = 0$
- Comparison result [Eschenburg, Heintze 90]:

$$\dot{\tilde{B}} + \tilde{B}^2 + \tilde{R} = 0$$
 and $\begin{array}{c} R \geq \tilde{R} ext{ on } I \\ B(0) \leq \tilde{B}(0) \end{array} \Rightarrow B \leq \tilde{B} ext{ on } I \cap [0,\infty) \end{array}$

• Choosing \tilde{R} and $\tilde{B}(t_0)$

• (1) suggests
$$\tilde{R} := \operatorname{diag}(c, -C, \ldots, -C), I = [-r, r]$$

• reasonably
$$B(0) := f(T, \delta, r) \cdot id$$

$$\implies \tilde{B} = \frac{1}{d} \operatorname{diag}(H_{c,f}, \dots, H_{-C,f}) \text{ (diagonal & explicit)}$$
$$\implies \operatorname{eigenvalue} \beta_{\min}(t) \le H_{c,f}(t) < H_{c,f}(\frac{r}{2}) < 0 \text{ on } [\frac{r}{2}, r]$$

• Feed into the shear term $\operatorname{tr}(\sigma^2)$ in the Raychaudhuri eq.: Integrating from $\frac{r}{2}$ to r contradicts boxing lemma for $T > T_0(r, c)$ and $\delta < \delta_0(r, c) \Rightarrow$ conjugate points in [-T, T].

The bound T_0 depends only on c, r not on g_{ε} !

Going back to $g \in C^{1,1}$

Shown so far:

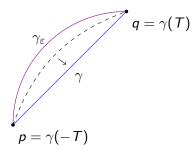
- $\check{g}_{arepsilon}\in C^{\infty}$ close to $g\in C^{1,1}$ which satisfies genericity and SEC
- γ_{ε} causal \check{g}_{ε} -geodesics close to γ causal g-geodesic
- $\Rightarrow \gamma_{\varepsilon}$ have conj. pts. if too long (longer than bd. uniform in ε)

Want to show: γ is not g-maximizing

Theorem (timelike case)

Let $g \in C^{1,1}$ be a globally hyperbolic Lorentzian metric on M satisfying genericity and SEC. Then any timelike geodesic γ is not globally maximising.

Proof



- Proof by contradiction, assume $\gamma : \mathbb{R} \to M$ is maximizing and satisfies genericity at t = 0
- Choose T > T₀(c, r), set
 p := γ(−T), q := γ(T)
- g glob. hyp. $\Rightarrow \check{g}_{\varepsilon}$ glob. hyp.
- $\exists \check{g}_{\varepsilon}$ -maximizing geodesics $\gamma_{\epsilon}: I_{\varepsilon} \to M$ from p to q
- Extract a convergent subsequence
- Limit must equal γ (else two distinct g-maximizing curves)
- But then $I_{\varepsilon} \rightarrow [-T, T]$, contradicting γ_{ϵ} being \check{g}_{ε} -maximizing

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Further issues (mainly swept under the carpet)

The null case of the previous theorem

- We cannot use global hyperbolicity
- To produce long approximating null geodesics we need to rule out that they are closed null geodesics for g by hand
- \rightsquigarrow in the theorem we have to suppose the spacetime to be causal rather than chronological (as in the classical case)
- The initial conditions:
- (a) the hypersurface case simply rests on $C^{1,1}$ -causality
- (b) We extend the trapped (2D-)surface case to C⁰-submanifolds of arbitrary codimensions generalising a condition by [Galloway, Senovilla 2010] using it in the support sense.
- (c) The trapped point condition also needs to be formulated in the support sense using (b).

The Hawking–Penrose theorem in $C^{1,1}$

Theorem[Hawking, Penrose 1970] [Graf, Grant,Kunzinger, S. 2017]

- A C^2 $C^{1,1}$ -spacetime (M,g) is causally incomplete if
- (i a) (SEC) $\operatorname{Ric}(X, X) \ge 0$ a.e. for every causal vector XLip. local vector field X
- (i b) $C^{1,1}$ -genericity holds
 - (ii) (M, g) is chronological causal
- (iii) *M* contains one of the following
 - (a) a compact achronal set A without edge
 - (b) a trapped C^0 -surface S in the support sense
 - (c) a trapped point in the support sense

(d) a trapped C^0 -submanifold of co-dimension 2 < m < nsatisfying the Galloway-Senovilla condition in the support sense

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Thanks for your attention